

# The nonlinear $\sigma$ -model in two dimensions

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Quantum field-theory of low dimensional systems

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The classical Heisenberg model is the n = 3 case of the n-vector model

$$H = -J \sum_{\langle i,j \rangle} \mathbf{S}_i \cdot \mathbf{S}_j$$
$$J > 0, \ \mathbf{S}_i \in \mathbb{R}^3, |\mathbf{S}_i| = 1$$

$$Z = \sum_{\{\mathbf{S}_i\}} e^{-\beta H}$$
$$= \sum_{\{\mathbf{S}_i\}} e^{\tilde{g} \sum_{\langle i,j \rangle} \mathbf{S}_i \cdot \mathbf{S}_j}, \ \tilde{g} = J\beta$$





■ All lattice spins aligned at T = 0



 All lattice spins aligned at T = 0  Lattice spins randomly oriented at high temperatures







Correlated groups of spins as low temperatures are approached





## $\bullet \xi > a$







•  $a \rightarrow 0$ , lattice  $\rightarrow$  continuum





• For low temperatures  $\frac{\xi}{a} \to \infty$  if  $a \to 0$ 





$$\mathbf{S}_{j} \approx \mathbf{S}_{i} + \partial_{\mu} \mathbf{S}_{i} a_{\mu} + \frac{1}{2} \partial_{\mu} \partial_{\nu} \mathbf{S}_{i} a_{\mu} a_{\nu}$$



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$$\mathbf{S}_{j} \approx \mathbf{S}_{i} + \partial_{\mu} \mathbf{S}_{i} a_{\mu} + \frac{1}{2} \partial_{\mu} \partial_{\nu} \mathbf{S}_{i} a_{\mu} a_{\nu}$$

$$\begin{split} \mathbf{S}_{j} \cdot \mathbf{S}_{i} &\approx \mathbf{S}_{i} \cdot \mathbf{S}_{i} + \mathbf{S}_{i} \partial_{\mu} \mathbf{S}_{i} a_{\mu} + \frac{1}{2} \mathbf{S}_{i} \partial_{\mu} \partial_{\nu} \mathbf{S}_{i} a_{\mu} a_{\nu} \\ &= 1 + \frac{1}{2} \partial_{\mu} \left( \mathbf{S}_{i} \cdot \mathbf{S}_{i} \right) a_{\mu} + \frac{1}{2} \mathbf{S}_{i} \partial_{\mu} \partial_{\mu} \mathbf{S}_{i} a^{2} \\ &= 1 + \frac{1}{2} \mathbf{S}_{i} \left( \partial_{\mu} \right)^{2} \mathbf{S}_{i} a^{2} \end{split}$$



$$\sum_{\langle i,j \rangle} \mathbf{S}_i \mathbf{S}_j \approx 2 \cdot \sum_i \frac{1}{2} \mathbf{S}_i (\partial_\mu)^2 \mathbf{S}_i a^2$$
$$= \frac{1}{a^d} \int \mathbf{S} (\partial_\mu)^2 \mathbf{S} a^2 d^d x$$
$$= \frac{-1}{a^{d-2}} \int \partial_\mu \mathbf{S} \ \partial_\mu \mathbf{S} \ d^d x$$
$$\sum e^{\tilde{g} \sum_{\langle i,j \rangle} \mathbf{S}_i \cdot \mathbf{S}_j} \rightarrow \int \mathcal{D} \mathcal{S} \ \delta \left( \mathbf{S}^2 - 1 \right) \ e^{\frac{-\tilde{g}}{a^{d-2}}} \ \int \partial_\mu \mathbf{S} \ \partial_\mu \mathbf{S} \ d^d x$$

$$Z = \sum_{\{\mathbf{S}_i\}} e^{j \, \mathcal{L}_{\langle i,j \rangle} \, \mathcal{S}_i \, \mathcal{S}_j} \to \int \mathcal{DS} \, \delta \left(\mathbf{S}^2 - 1\right) \, e^{ad-2} \, j^{|\mathbf{S}_\mu| \mathcal{S}_i \, \mathbf{S}_\mu| \mathcal{S}_i \, \mathbf{S}_\mu}$$

Continuum limit  $\rightarrow$  nonlinear  $\sigma$ -model

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$$S = \frac{1}{g} \int \partial_{\mu} \mathbf{S} \ \partial_{\mu} \mathbf{S} \ d^{d}x \quad \frac{1}{g} = \frac{\tilde{g}}{a^{d-2}}$$









Divide the solid into blocks of 2 ×2 squares

$$Z = \sum_{\{\sigma\}} \exp\left(\sum_{\langle i,j \rangle} J_{ij} \sigma_i \sigma_j\right)$$
$$\{\sigma_1, \sigma_2, \sigma_3, \sigma_4\} \to \sigma_1'$$

$$Z = \sum_{\{\sigma'\}} \exp\left(\sum_{\langle i,j \rangle} J'_{ij} \sigma'_i \sigma'_j\right)$$
$$\{\sigma'_1, \sigma'_2, \sigma'_3, \sigma'_4\} \to \sigma''_1$$

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Divide the solid into blocks of 2 ×2 squares

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$$\begin{split} Z = \sum_{\{\sigma'\}} \exp\left(\sum_{\langle i,j\rangle} J'_{ij} \sigma'_i \sigma'_j\right) \\ \left\{\sigma'_1, \sigma'_2, \sigma'_3, \sigma'_4\right\} \to \sigma''_1 \end{split}$$

....



The change in the parameters is implemented by a certain β-function:

$$\kappa \frac{\partial g}{\partial \kappa} = \beta(g)$$

- $\beta$ -function  $\rightarrow$  RG flow. The values of g under the flow are called running couplings
- The possible macroscopic states of a system, at a large scale, are given by a set of fixed points



RG-flow equation

$$\kappa \ \frac{\partial g}{\partial \kappa} = \beta(g)$$

- small  $\kappa \to \text{flow in the}$ infra-red
- no fixed point at finite temperatures for  $d \le 2$
- for d>2 theory has a fixed point at  $T_c \rightarrow$  phase transition





$$\mathcal{L} = \frac{1}{2} \partial_{\mu} \mathbf{\Phi} \cdot \partial^{\mu} \mathbf{\Phi} - U(\mathbf{\Phi}(x,t)), \quad \mathbf{\Phi}(\mathbf{x},\mathbf{t}) = [\Phi_i(x,t); i = 1, ..., N]$$

$$\begin{split} (\delta S)_{\Phi} &= \delta \int \left[ \frac{1}{2} \partial_{\mu} \mathbf{\Phi} \cdot \partial^{\mu} \mathbf{\Phi} - U \left( \mathbf{\Phi}(x, t) \right) \right] d^{d}x = 0 \\ &= \int \left[ -\partial_{\mu} \partial^{\mu} \mathbf{\Phi} - \frac{dU}{d\mathbf{\Phi}} \right] \delta \mathbf{\Phi} d^{d}x = 0 \end{split}$$

$$\Rightarrow \partial_{\mu}\partial^{\mu}\Phi + \frac{dU}{d\Phi} = \left(\partial_{t}^{2} - \nabla^{2}\right)\Phi + \frac{dU}{d\Phi} = 0$$



- $\blacksquare~U\left( {\bf \Phi}(x,t) \right)$  ist assumed to be non-negative and only vanishes at its absolute minima  $U({\bf \Phi}) \geq 0$
- A static solution  $\mathbf{\Phi}(x)$  obeys

$$\nabla^2 \Phi = \frac{dU}{d\Phi} \tag{1}$$

• (1) extremum condition for  $W[\Phi] \rightarrow \delta W[\Phi] = 0$ 

$$\begin{split} W[\Phi] &= \int \left[ \frac{1}{2} \nabla_i \Phi \cdot \nabla_i \Phi + U\left( \Phi(x) \right) \right] d^d x \\ &= V_1[\Phi] + V_2[\Phi] \end{split}$$



• Cosider one parameter family of solutions of a static solution  $\Phi_1(x)$ 

$$\begin{split} \boldsymbol{\Phi}_{\lambda} &= \boldsymbol{\Phi}_{1}(\lambda x) \\ \Rightarrow W[\Phi_{\lambda}] &= V_{1}[\Phi_{\lambda}] + V_{2}[\Phi_{\lambda}] \\ &= \lambda^{2-d}V_{1}[\Phi] + \lambda^{-d}V_{2}[\Phi] \end{split}$$

•  $\Phi_1(x)$  is an extremum of  $W[\Phi] \Rightarrow \frac{d}{d\lambda} W[\Phi_{\lambda}] = 0$  at  $\lambda = 1$ 

$$(2-d) V_1[\Phi_1] = d V_2[\Phi_1]$$

• There is no non-trivial static space-dependent solution for  $d \ge 3$ .



The model consists of 3 scalar fields

$$\mathbf{\Phi}(\mathbf{x},t) = \{\Phi_1(\mathbf{x},t), \Phi_2(\mathbf{x},t), \Phi_3(\mathbf{x},t)\}$$

The dynamics of the system is given by a Lagrangian density

$$\mathcal{L} = \frac{1}{2} \sum_{\mu} \sum_{a} \left( \partial_{\mu} \Phi_{a} \partial^{\mu} \Phi_{a} \right) = \frac{1}{2} \left( \partial_{\mu} \Phi \right) \cdot \left( \partial^{\mu} \Phi \right)$$

The system is subject to a constraint

$$\sum_{a} \Phi_a(\mathbf{x}, t)^2 = \mathbf{\Phi} \cdot \mathbf{\Phi} = 1$$



Constraint is imposed via a Lagrange multiplier

$$S[\mathbf{\Phi}] = \int d^d x \int dt \left[ \frac{1}{2} \left( \partial_\mu \mathbf{\Phi} \right) \cdot \left( \partial^\mu \mathbf{\Phi} \right) + \lambda(\mathbf{x}, t) \left( \mathbf{\Phi} \cdot \mathbf{\Phi} - 1 \right) \right]$$

Obtained field equations from the Lagrangian

$$\partial_{\mu}\partial^{\mu}\Phi + \lambda\Phi = (\Box + \lambda)\Phi = 0$$

Lagrange mutliplier is eliminated by the given constraint

$$\lambda(\mathbf{x},t) = \lambda \boldsymbol{\Phi} \cdot \boldsymbol{\Phi} = -\boldsymbol{\Phi} \Box \boldsymbol{\Phi}$$



Field equations for static solutions in 2 dimensions

$$\nabla^2 \mathbf{\Phi} - \left( \mathbf{\Phi} \cdot \nabla^2 \mathbf{\Phi} \right) \mathbf{\Phi} = 0$$

 $\blacksquare$  Constraint  $\Rightarrow$  field configurations can be classified into homotopy sectors



### Definition

Let X, Y be topological spaces and  $f, g: X \to Y$  continuous maps.

A homotopy from f to g is a continuous function  $H: X \times [0,1] \to Y$  satisfying H(x,0) = f(x) and H(x,1) = g(x), for all  $x \in X$ .

If such a homotopy exists , we say that f is homotopic to g, and denote this by  $f\simeq g.$ 









 $S^1 \to S^1$ 



Energy of a static solution

$$E = \frac{1}{2} \int (\partial_{\mu} \mathbf{\Phi}) \cdot (\partial_{\mu} \mathbf{\Phi}) d^{2}x$$

Solutions for E = 0

$$\Rightarrow \partial_{\mu} \mathbf{\Phi}(\mathbf{x}) = 0 \quad \forall \mathbf{x} \Rightarrow \mathbf{\Phi}(\mathbf{x}) = \mathbf{\Phi}^{(0)}$$

• O(3) symmetry  $\Rightarrow$  continuous family of degenerate classical minima



$$\frac{1}{2}\int\left(\partial_{\mu}\Phi\right)\cdot\left(\partial_{\mu}\Phi\right)d^{2}x>0$$
 but finite

• Using polar coordinates  $(r, \phi)$ , finite energy solutions in  $\mathbb{R}^2$  must satisfy the following conditions:

$$\begin{split} r \, ||\mathbf{grad} \mathbf{\Phi}|| &\to 0 \quad \text{as } r \to \infty \\ \lim_{n \to \infty} \mathbf{\Phi}(x) &= \mathbf{\Phi}^{(0)} \quad \text{since} \quad \partial_2 \mathbf{\Phi} = \frac{1}{r} \frac{\partial \mathbf{\Phi}}{\partial \phi} \end{split}$$







Since  $\Phi(x)$  approaches the same value  $\Phi^{(0)}$  at all points in infinity, the plane in  $\mathbb{R}^2$  can be compactified into a spherical surface  $S_2^{(phy)}$ 



The circle at infinity is reduced to the northpole of the sphere



• Internal space  $S_2^{(\text{int})}$  and coordinate space  $S_2^{(\text{phy})}$  are both spherical surfaces  $\Rightarrow \mathbf{\Phi}: S_2^{(\text{phy})} \rightarrow S_2^{(\text{int})}$ 



 All non-singular mappings of S<sub>2</sub> into S<sub>2</sub> can be classified into homotopy sectors.



## Q can be written as an integral

$$Q = \frac{1}{8\pi} \int \epsilon_{\mu\nu} \mathbf{\Phi} \cdot \left(\partial_{\mu} \mathbf{\Phi} \times \partial_{\nu} \mathbf{\Phi}\right) d^2 x$$

•  $S_2^{(\text{int})}$  can be described by  $\{\xi_1, \xi_2\} \rightarrow dS_a^{(\text{int})} = \frac{1}{2} \epsilon_{rs} \epsilon_{abc} \frac{\partial \Phi_b}{\partial \xi_r} \frac{\partial \Phi_c}{\partial \xi_s} d^2 \xi$ 

$$\begin{split} Q &= \frac{1}{8\pi} \int \epsilon_{\mu\nu} \epsilon_{abc} \Phi_a \frac{\partial \Phi_b}{\partial x^{\mu}} \frac{\partial \Phi_c}{\partial x^{\nu}} d^2 x \\ &= \frac{1}{8\pi} \int \epsilon_{\mu\nu} \epsilon_{abc} \Phi_a \frac{\partial \Phi_b}{\partial \xi_r} \frac{\partial \xi_r}{\partial x_{\mu}} \frac{\partial \Phi_c}{\partial \xi_s} \frac{\partial \xi_s}{\partial x_{\nu}} d^2 x \\ Q &= \frac{1}{8\pi} \int \epsilon_{rs} \epsilon_{abc} \Phi_a \frac{\partial \Phi_b}{\partial \xi_r} \frac{\partial \Phi_c}{\partial \xi_s} d^2 \xi \quad \text{as} \quad \epsilon_{rs} d^2 \xi = \epsilon_{\mu\nu} \frac{\partial \xi_r}{\partial x_{\mu}} \frac{\partial \xi_s}{\partial x_{\nu}} d^2 x \end{split}$$



$$dS_a^{(\text{int})} = \frac{1}{2} \epsilon_{rs} \epsilon_{abc} \frac{\partial \Phi_b}{\partial \xi_r} \frac{\partial \Phi_c}{\partial \xi_s} d^2 \xi$$
(2)  
$$Q = \frac{1}{8\pi} \int \epsilon_{rs} \epsilon_{abc} \Phi_a \frac{\partial \Phi_b}{\partial \xi_r} \frac{\partial \Phi_c}{\partial \xi_s} d^2 \xi$$
(3)

Inserting (2) in (3) gives

$$Q = \frac{1}{4\pi} \int dS_a^{(\text{int})} \Phi_a = \frac{1}{4\pi} \int dS^{(\text{int})}$$

• Q gives the number of times the internal sphere is traversed as the coordinate space  $\mathbb{R}^2$  (which is compacted into  $S_2^{(phy)}$ ) is spanned



 Homotopy classification is valid for any static field configuration with finite energy → not necessarily a solution of the field equation
Using the following identity

$$\int \left[ (\partial_{\mu} \Phi \pm \epsilon_{\mu\nu} \Phi \times \partial_{\nu} \Phi) \cdot (\partial_{\mu} \Phi \pm \epsilon_{\mu\sigma} \Phi \times \partial_{\sigma} \Phi) \right] d^2 x \ge 0 \quad (4)$$

Expanding (4) and using the constraint  ${f \Phi}\cdot{f \Phi}=1$  yields

$$2\int d^2x \left(\partial_\mu \Phi\right) \cdot \left(\partial_\mu \Phi\right) \ge \pm 2\int d^2x \epsilon_{\mu\nu} \Phi \cdot \left(\partial_\mu \Phi \times \partial_\nu \Phi\right)$$
$$E \ge 4\pi \left|Q\right|$$



In any given sector Q the energy is minimised when  $E=4\pi \left|Q\right|$  is satisfied. This happens if and only if

$$\partial_{\mu} \Phi = \pm \epsilon_{\mu\nu} \Phi \times (\partial_{\nu} \Phi)$$
 is satisfied (5)

• (5) can be further simplified by using a stereographic projection  $\mathbf{\Phi} = {\Phi_1, \Phi_2, \Phi_3} \rightarrow {\omega_1, \omega_2}$ 

$$\omega_1 = \frac{2\Phi_1}{1 - \Phi_3} \quad \omega_2 = \frac{2\Phi_2}{1 - \Phi_3}$$



• The stereographic projection  $\{\Phi_1, \Phi_2, \Phi_3\} \rightarrow \{\omega_1, \omega_2\}$  gives us the Cauchy-Riemann equations

$$\frac{\partial \omega_1}{\partial x_1} = \pm \frac{\partial \omega_2}{\partial x_2}$$
$$\frac{\partial \omega_1}{\partial x_2} = \mp \frac{\partial \omega_2}{\partial x_1}$$

• A prototype solution for an arbitrary positive Q is given by

$$\omega(z) = \frac{\left[(z - z_0)\right]^n}{\lambda^n}, n \in \mathbb{N}, \lambda \in \mathbb{R}, z \in \mathbb{C}$$



$$\omega_1 = \frac{2\Phi_1}{1 - \Phi_3} \quad \omega_2 = \frac{2\Phi_2}{1 - \Phi_3} \quad \omega(z) = \left(\frac{z - z_0}{\lambda}\right)^n$$

• Solution for n = 1 and  $z_0 = 0$ 

$$|\omega|^2 = \omega_1^2 + \omega_2^2 = \frac{r^2}{\lambda^2} \qquad \Phi_1 = \omega_1 \frac{(1 - \Phi_3)}{2} \qquad \Phi_2 = \omega_2 \frac{(1 - \Phi_3)}{2}$$

• Using the constraint 
$$\Phi_1^2 + \Phi_2^2 + \Phi_3^2 = 1 \rightarrow \Phi_3 = \frac{\frac{r^2}{\lambda^2} - 1}{\frac{r^2}{\lambda^2} + 1}$$

I

$$\mathbf{\Phi} = \frac{1}{1 + \left(\frac{r}{2\lambda}\right)^2} \left\{ \frac{r\cos(\phi)}{\lambda}, \frac{r\sin(\phi)}{\lambda}, \left(\frac{r}{2\lambda}\right)^2 - 1 \right\}$$







• Continuum limit  $\rightarrow$  nonlinear  $\sigma$ -model

$$S = \frac{1}{g} \int \partial_{\mu} \Phi_a \partial_{\mu} \Phi_a d^d x$$

- Constraint  $\mathbf{\Phi} \cdot \mathbf{\Phi} = 1 \rightarrow \nabla^2 \mathbf{\Phi} \left(\mathbf{\Phi} \cdot \nabla^2 \mathbf{\Phi}\right) \mathbf{\Phi} = 0$
- Finite energy solutions:  $\mathbb{R}^2 \to S^2$
- Homotopy classification for static field configurations

$$Q = \frac{1}{8\pi} \int \epsilon_{rs} \epsilon_{abc} \Phi_a \frac{\partial \Phi_a}{\partial \xi_r} \frac{\partial \Phi_c}{\partial \xi_s} d^2 \xi$$
$$E \ge 4\pi |Q|$$



 $\blacksquare$  Lower bound for the energy  $E \geq 4\pi \left| Q \right| \Rightarrow$  first-order differential equation

$$\partial_{\mu} \Phi = \pm \epsilon_{\mu\nu} \Phi \times (\partial_{\nu} \Phi) \tag{6}$$

• Using a stereographic projection  $S_2^{(int)} \to \mathbb{R}^2 \Rightarrow \omega = \omega_1 + i\omega_2$ 

$$\frac{\partial \omega_1}{\partial x_1} = \pm \frac{\partial \omega_2}{\partial x_2} \quad \frac{\partial \omega_1}{\partial x_2} = \mp \frac{\partial \omega_2}{\partial x_1}$$

 $\blacksquare$  Solution to Cauchy-Riemann equations is well known  $\rightarrow$  Skyrmion

$$\mathbf{\Phi} = \frac{1}{1 + \left(\frac{r}{2\lambda}\right)^2} \left\{ \frac{r\cos(\phi)}{\lambda}, \frac{r\sin(\phi)}{\lambda}, \left(\frac{r}{2\lambda}\right)^2 - 1 \right\}$$



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